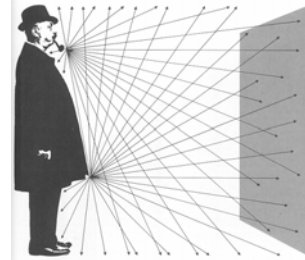


The Light Field

Concepts

- Light field = radiance function on rays
- Conservation of radiance
- Throughput and counting rays
- Measurement equation
- Area sources and irradiance



From London and Upton

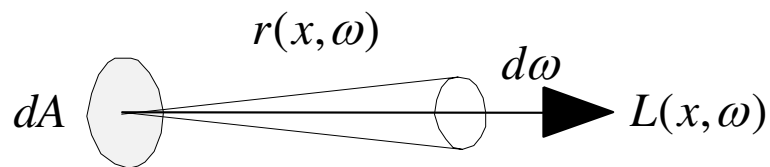
CS348B Lecture 5

Pat Hanrahan, 2006

Light Field = Radiance(Ray)

Field Radiance

Definition: The field *radiance* (*luminance*) at a point in space in a given direction is the power per unit solid angle per unit area perpendicular to the direction



Radiance is the quantity associated with a ray

CS348B Lecture 5

Pat Hanrahan, 2006

Light Probe \Rightarrow Environment Map

$L(x, y, z, \theta, \phi)$

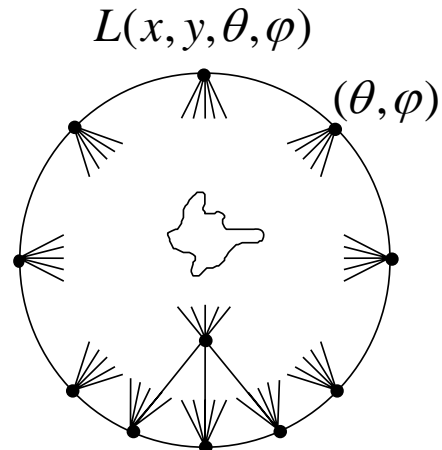


Miller and Hoffman, 1984

CS348B Lecture 5

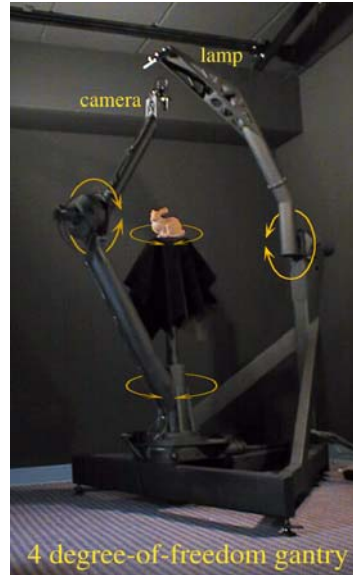
Pat Hanrahan, 2006

Spherical Gantry \Rightarrow 4D Light Field



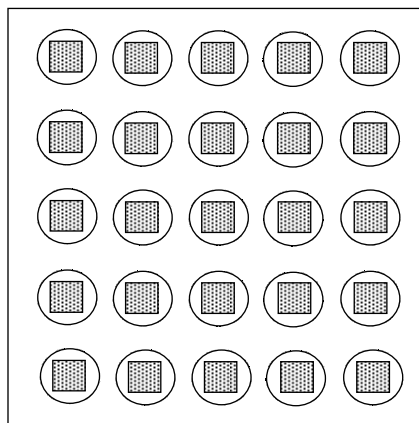
Capture all the light leaving
an object - like a hologram

CS348B Lecture 5

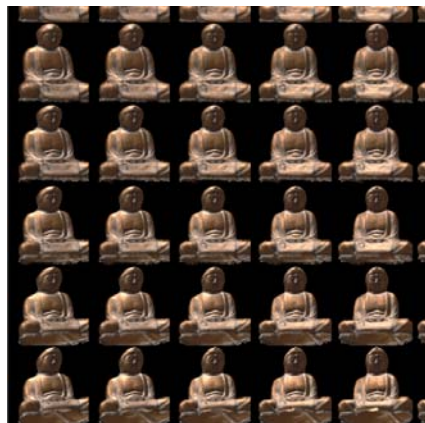


Pat Hanrahan, 2006

Two-Plane Light Field



2D Array of Cameras



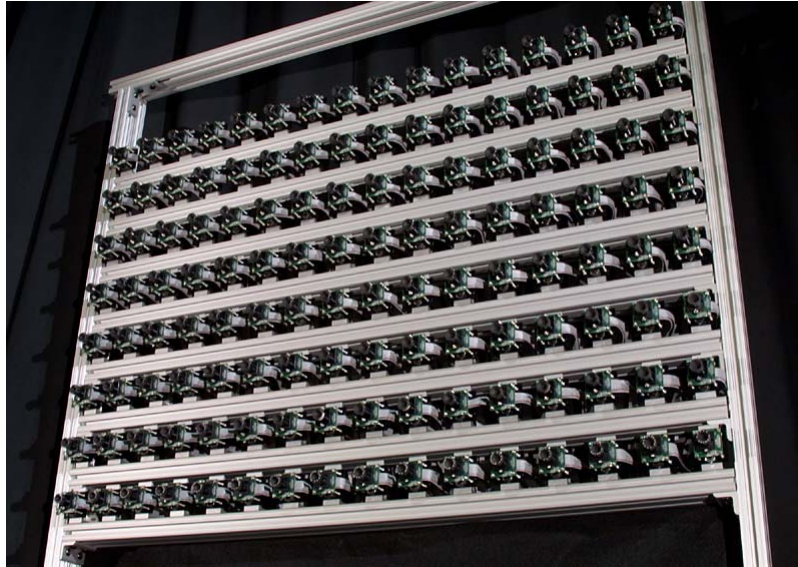
2D Array of Images

$$L(u, v, s, t)$$

CS348B Lecture 5

Pat Hanrahan, 2006

Multi-Camera Array \Rightarrow Light Field



CS348B Lecture 5

Pat Hanrahan, 2006

Properties of Radiance

Properties of Radiance

1. **Fundamental field quantity that characterizes the distribution of light in an environment.**
 - ∴ Radiance is a function on rays
 - ∴ All other field quantities are derived from it
2. **Radiance invariant along a ray.**
 - ∴ 5D ray space reduces to 4D
3. **Response of a sensor proportional to radiance.**

CS348B Lecture 5

Pat Hanrahan, 2006

1st Law: Conservation of Radiance

The radiance in the direction of a light ray remains constant as the ray propagates

$$d^2\Phi_1 = d^2\Phi_2$$

$$d^2\Phi_1 = L_1 d\omega_1 dA_1$$

$$d^2\Phi_2 = L_2 d\omega_2 dA_2$$

$$d\omega_1 dA_1 = \frac{dA_1 dA_2}{r^2} = d\omega_2 dA_2$$

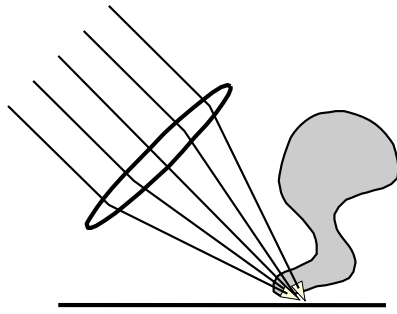
$$\therefore L_1 = L_2$$

CS348B Lecture 5

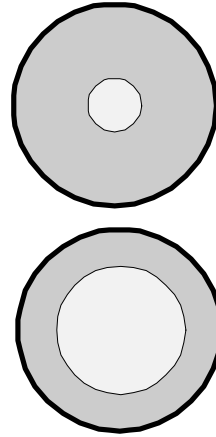
Pat Hanrahan, 2006

Quiz

Does radiance increase under a magnifying glass?



No!!



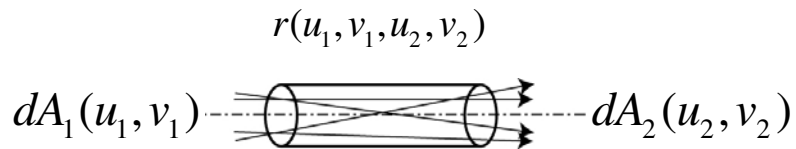
CS348B Lecture 5

Pat Hanrahan, 2006

Measuring Rays = Throughput

Throughput Counts Rays

Define an infinitesimal beam as the set of rays intersecting two infinitesimal surface elements



The differential throughput is the size of the beam

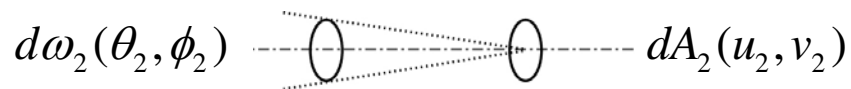
$$d^2T = \frac{dA_1 dA_2}{|x_1 - x_2|^2}$$

CS348B Lecture 5

Pat Hanrahan, 2006

Parameterizing Rays

Parameterize rays wrt to receiver $r(u_2, v_2, \theta_2, \phi_2)$



$$d^2T = \frac{dA_1}{|x_1 - x_2|^2} dA_2 = d\omega_2 dA_2$$

CS348B Lecture 5

Pat Hanrahan, 2006

Parameterizing Rays

Parameterize rays wrt to source $r(u_1, v_1, \theta_1, \phi_1)$



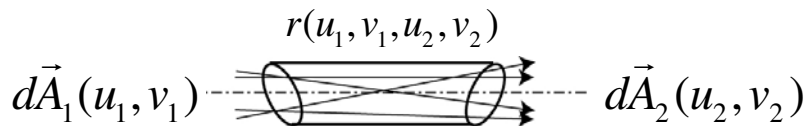
$$d^2T = dA_1 \frac{dA_2}{|x_1 - x_2|^2} = dA_1 d\omega_1$$

CS348B Lecture 5

Pat Hanrahan, 2006

Parameterizing Rays

Tilting the surfaces reparameterizes the rays



$$\begin{aligned} d^2T &= \frac{\cos \theta_1 \cos \theta_2}{|x_1 - x_2|^2} dA_1 dA_2 \\ &= d\vec{\omega}_1 \cdot d\vec{A}_1 \\ &= d\vec{\omega}_2 \cdot d\vec{A}_2 \end{aligned}$$

CS348B Lecture 5

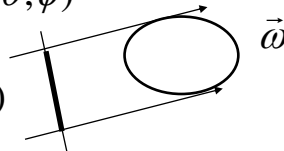
Pat Hanrahan, 2006

Parameterizing Rays: $S^2 \times R^2$

Parameterize rays by $r(x, y, \theta, \phi)$

Projected area

$$\tilde{A}(\vec{\omega})$$



Measuring the number or rays that hit a shape

$$\begin{aligned} T &= \int_{S^2} d\omega(\theta, \phi) \int_{R^2} dA(x, y) \\ &= \int_{S^2} \tilde{A}(\theta, \phi) d\omega(\theta, \phi) \\ &= 4\pi \bar{\tilde{A}} \end{aligned}$$

Sphere:

$$T = 4\pi \bar{\tilde{A}} = 4\pi^2 R^2$$

CS348B Lecture 5

Pat Hanrahan, 2006

Parameterizing Rays: $M^2 \times S^2$

Parameterize rays by $r(u, v, \theta, \phi)$

$$T = \underbrace{\left[\int_{M^2} dA(u, v) \right]}_S \underbrace{\left[\int_{H^2(\vec{N})} \cos \theta d\omega(\theta, \phi) \right]}_\pi$$

Sphere: $T = \pi S = 4\pi^2 R^2$

Crofton's Theorem: $4\pi \bar{\tilde{A}} = \pi S \Rightarrow \bar{\tilde{A}} = \frac{S}{4}$

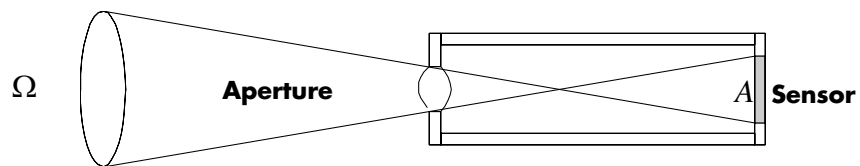
CS348B Lecture 5

Pat Hanrahan, 2006

The Measurement Equation

Radiance: 2nd Law

The response of a sensor is proportional to the radiance of the surface visible to the sensor.



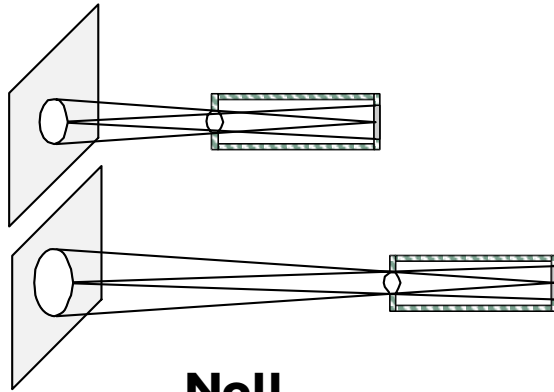
$$R = \int \int_{A \Omega} L d\omega dA = \bar{L} T \quad T = \int \int_{A \Omega} d\omega dA$$

L is what should be computed and displayed.

T quantifies the gathering power of the device; the higher the throughput the greater the amount of light gathered

Quiz

Does the brightness that a wall appears to the sensor depend on the distance?



CS348B Lecture 5

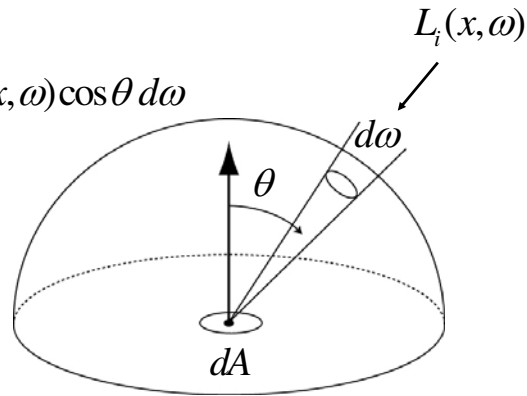
Pat Hanrahan, 2006

Irradiance from a Uniform Area Source

Irradiance from the Environment

$$d^2\Phi_i(x, \omega) = L_i(x, \omega) \cos \theta dA d\omega$$

$$dE(x, \omega) \equiv \frac{d^2\Phi}{dA} = L_i(x, \omega) \cos \theta d\omega$$



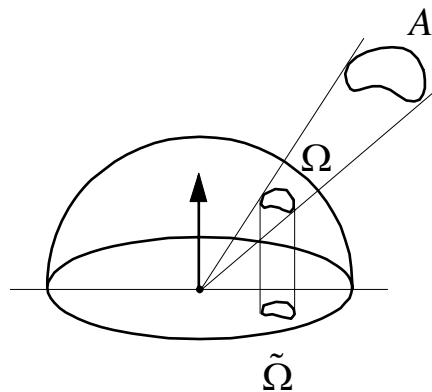
$$E(x) = \int_{H^2} L_i(x, \omega) \cos \theta d\omega$$

CS348B Lecture 5

Pat Hanrahan, 2006

Uniform Area Source

$$\begin{aligned} E(x) &= \int_{H^2} L \cos \theta d\omega \\ &= L \int_{\Omega} \cos \theta d\omega \\ &= L \tilde{\Omega} \end{aligned}$$

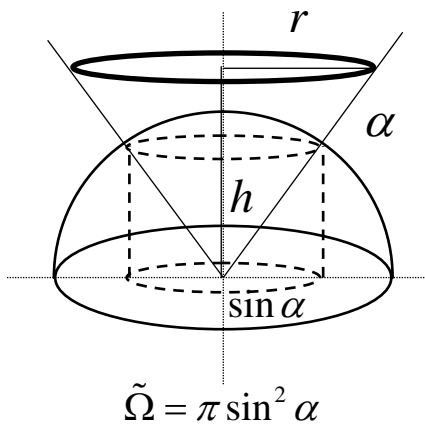


CS348B Lecture 5

Pat Hanrahan, 2006

Uniform Disk Source

Geometric Derivation



Algebraic Derivation

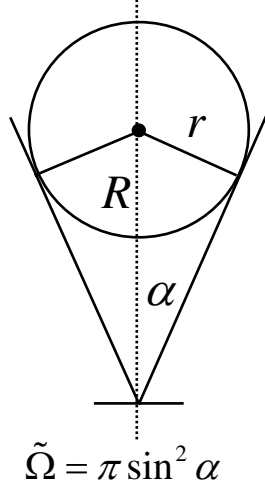
$$\begin{aligned}\tilde{\Omega} &= \int_1^{\cos \alpha} \int_0^{2\pi} \cos \theta \, d\phi \, d \cos \theta \\ &= 2\pi \frac{\cos^2 \theta}{2} \Big|_1^{\cos \alpha} \\ &= \pi \sin^2 \alpha \\ &= \pi \frac{r^2}{r^2 + h^2}\end{aligned}$$

CS348B Lecture 5

Pat Hanrahan, 2006

Spherical Source

Geometric Derivation



Algebraic Derivation

$$\begin{aligned}\tilde{\Omega} &= \int \cos \theta \, d\omega \\ &= \pi \sin^2 \alpha \\ &= \pi \frac{r^2}{R^2}\end{aligned}$$

CS348B Lecture 5

Pat Hanrahan, 2006

The Sun

Solar constant (normal incidence at zenith)

Irradiance **1353 W/m²**

Illuminance **127,500 lm/m² = 127.5 kilolux**

Solar angle

$\alpha = .25 \text{ degrees} = .004 \text{ radians (half angle)}$

$\tilde{\Omega} = \pi \sin^2 \alpha \approx \pi \alpha^2 = 6 \times 10^{-5} \text{ steradians}$

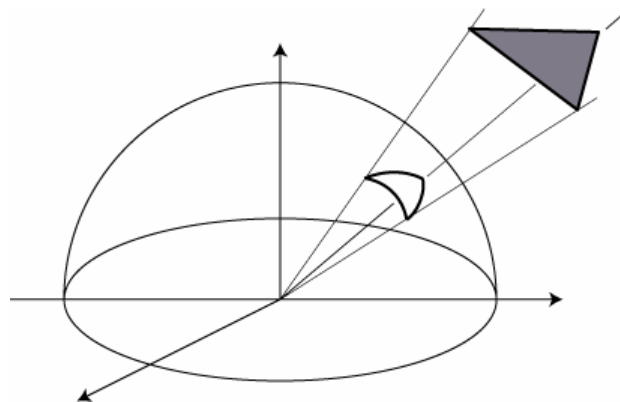
Solar radiance

$$L = \frac{E}{\tilde{\Omega}} = \frac{1.353 \times 10^3 \text{ W/m}^2}{6 \times 10^{-5} \text{ sr}} = 2.25 \times 10^7 \frac{\text{W}}{\text{m}^2 \cdot \text{sr}}$$

CS348B Lecture 5

Pat Hanrahan, 2006

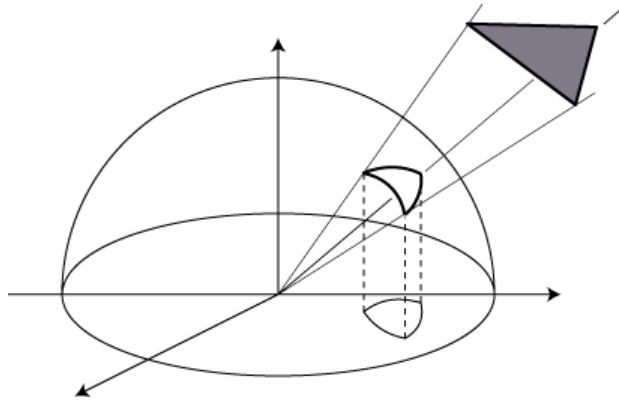
Polygonal Source



CS348B Lecture 5

Pat Hanrahan, 2006

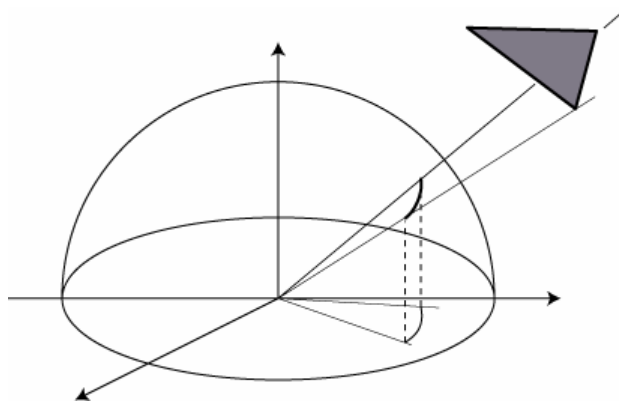
Polygonal Source



CS348B Lecture 5

Pat Hanrahan, 2006

Polygonal Source

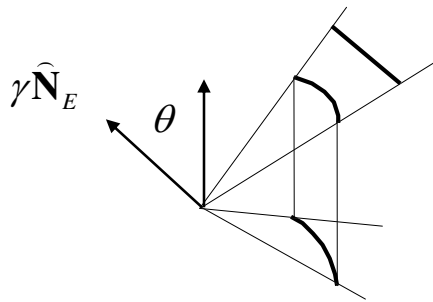


CS348B Lecture 5

Pat Hanrahan, 2006

Consider 1 Edge

γ_i Area of sector

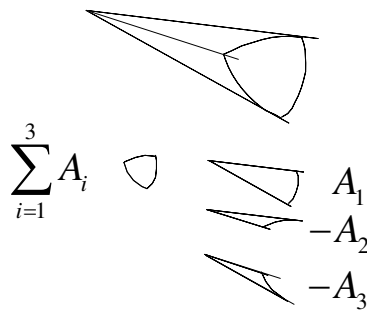


$$A = \gamma \cos \theta = \gamma \vec{N}_E \cdot \vec{N}$$

CS348B Lecture 5

Pat Hanrahan, 2006

Lambert's Formula



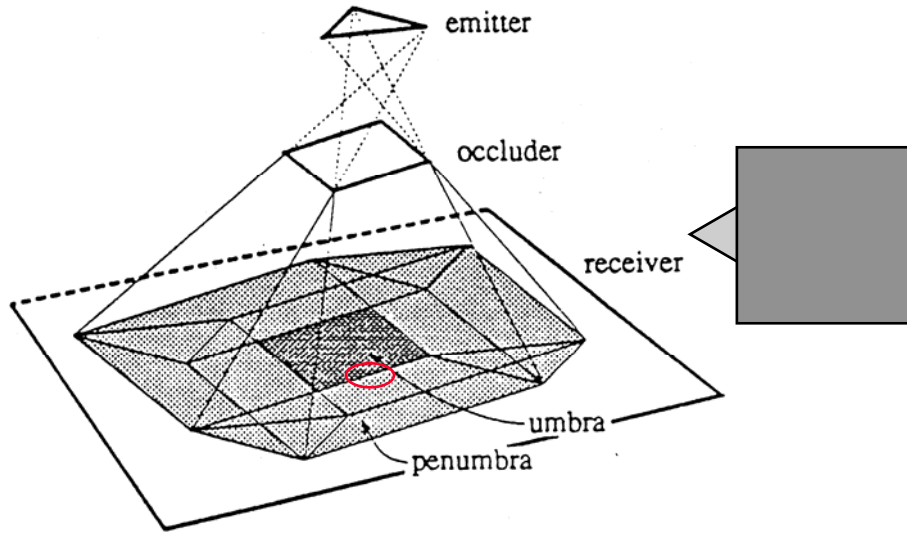
$$\sum_{i=1}^3 A_i$$

$$\sum_{i=1}^n A_i = \sum_{i=1}^n \gamma_i \vec{N}_i \cdot \vec{N}$$

CS348B Lecture 5

Pat Hanrahan, 2006

Penumbras and Umbras



CS348B Lecture 5

Pat Hanrahan, 2006